# Incorporation of ferric iron in CaSiO<sub>3</sub> perovskite at high pressure

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## ABSTRACT

Synthetic andradite (Ca<sub>3</sub>Fe<sub>2</sub>Si<sub>3</sub>O<sub>12</sub>) has been compressed to loading pressures >21 GPa and heated to ~1000°C by a YAG laser in a Diamond Anvil Cell (DAC). After quenching to room temperature, X-ray diffraction of the sample, still held at 21 GPa, showed that andradite had transformed to a cubic perovskite type polymorph with a = 3.460(4) Å. Upon decompression the perovskite phase transformed into an amorphous phase. The density of the perovskite polymorph (Ca<sub>3</sub>Fe<sub>2</sub>Si<sub>3</sub>O<sub>12</sub>) is ~13.6% greater than that of isochemical andradite at 21 GPa. Ferric iron replaces Ca<sup>2+</sup> and Si<sup>4+</sup> in the perovskite structure (Fe<sup>3+</sup> + Fe<sup>3+</sup> = Si<sup>4+</sup> + Ca<sup>2+</sup>), giving a formula unit: (Ca<sub>5</sub>Fe<sup>3+</sup>)(Si<sub>5</sub>Fe<sup>3+</sup>)O<sub>3</sub>. The new Fe<sup>3+</sup>-rich Ca-perovskite may provide new insight into the controls on the electrical conductivity of the lower mantle.

KEYWORDS: andradite, high pressure, X-ray diffraction, perovskite, lower mantle.

#### Introduction

BECAUSE of the special significance of ferric iron for the geophysics of solid earth interior, the presence of ferric iron in mantle minerals is of particular current interest. Evidence from in situ high-pressure Mössbauer spectra and from the components of diamond inclusions derived from the lower mantle shows that ferric iron can be incorporated into perovskite (Mg,Fe)SiO<sub>3</sub> and magnesiowüstite (Fe,Mg)O (Harris et al., 1997; McCammon et al., 1997). We have studied and radite  $(Ca_3Fe_2Si_3O_{12})$  by high-pressure X-ray diffraction to explore its stability at mantle conditions and to obtain a deeper understanding of the structural aspects of its high-pressure phase transitions, in the context of the geophysical properties of the Earth's interior. In this letter, the results of these high-pressure X-ray diffraction investigations are presented and discussed.

#### Experimental

#### High-pressure synthesis of andradite

 $CaCO_3$ ,  $Fe_2O_3$  and  $SiO_2$  were mixed in a molar ratio of 3:1:3 and ground in a ceramic mortar. The

initial mixture was held at  $1000^{\circ}$ C for 24 h to decarbonate the CaCO<sub>3</sub>. Thus, a starting material was obtained with a bulk composition near stoichiometric Ca<sub>3</sub>Fe<sub>2</sub>Si<sub>3</sub>O<sub>12</sub>, which was used for subsequent high-pressure synthesis.

High-pressure runs were carried out using a wedge-type cubic anvil apparatus (Hishinuma and Yagi, 1994). Temperatures were measured using a Pt-Pt13%Rh thermocouple. The pure andradite was formed at 1150°C and 3 GPa in 1 h runs. The unit-cell parameter was determined by conventional X-ray powder diffraction (monochromatic Cr-K $\alpha_1$  radiation) from at least 10 peaks at >100° 2 $\theta$ . The lattice parameter of the synthetic andradite was a = 12.068(2) Å (Ia3d), in good agreement with the pioneering results of Woodland and O'Neill (1995).

#### In situ high-pressure X-ray diffraction

High-pressure X-ray diffraction measurements were carried out at the synchrotron radiation source at the High Energy Physics Institute (KEK) using a modified Mao-Bell type Diamond Anvil Cell (DAC). The diamond culet was 450 µm in diameter. The powder sample was squeezed



Fig. 1. Room temperature and pressure X-ray diffraction pattern of a sample recovered from a high-P/T run (26.9 GPa with laser heating). The seven weak peaks indicated are attributed to and radite.

between two diamond anvils and compacted into a disk (150  $\mu$ m in diameter and 10  $\mu$ m thick) in a stainless steel gasket.

The pressure was calibrated using the R<sub>1</sub> fluorescence of ruby (Mao *et al.*, 1978). After the fine powdered sample had been compressed to 20–30 GPa in the DAC, it was heated to ~1000°C using a YAG laser. This process was carried out for 2 h to make sure the sample was heated sufficiently. The recovered sample, retained at high pressure within the DAC, was observed by optical microscopy and then examined by X-ray diffraction to identify the phases present under temperature-quenched conditions. Furthermore, monochromatic synchrotron radiation ( $\lambda = 0.4313$  Å,  $2\theta = 5-25^{\circ}$ ) was employed to chart any structural changes upon release of pressure.

#### **Results and discussion**

*In situ* high-pressure optical observation revealed that, before laser heating, the sample was transparent, while after laser heating it had become optically opaque. The change in colour was the first important indicator of the existence of a possible phase transformation. X-ray diffraction on the recovered sample at 0 GPa showed that andradite had transformed into an amorphous phase, showing several weak diffraction peaks upon a high background (Fig. 1).

The X-ray diffraction patterns obtained from the temperature-quenched sample at 1 atmosphere and 21 GPa are shown in Fig. 2. Eight intense lines in the high-pressure diffraction pattern could not be attributed to andradite. The positions of these peaks are similar to those of  $CaSiO_3$ perovskite (Liu and Ringwood, 1975; Yagi *et al.*, 1989), and all of them can readily be indexed in terms of the cubic perovskite structure (Table 1). At 21 GPa, the new phase of  $Ca_3Fe_2Si_3O_{12}$  shows some very strong peaks, and upon decompression this new phase was retained, but as the pressure approached 0 GPa its diffraction signal disappeared as it transformed into an amorphous phase. It is worth pointing out that andradite peaks are present in all patterns, coexisting with the cubic phase due to incomplete transformation.



FIG. 2. X-ray diffraction patterns of  $Ca_3Fe_2Si_3O_{12}$  at pressures of 1 atmosphere and 21 GPa. Indexed peaks refer to the cubic perovskite structure.



hkl	$d_{(\text{obs})}$ (Å)	$d_{(\text{cal})}$ (Å)	$d_{(o)}/d_{(c)} - 1$
100	3.468	3.460	0.002
110	2.462	2.447	0.006
111	1.995	1.998	-0.001
200	1.731	1.730	0.001
210	1.544	1.547	-0.002
211	1.422	1.413	0.007
220	1.223	1.223	-0.001
310	1.092	1.094	-0.002

TABLE 1. The Bragg positions of Ca-Fe-perovskite obtained *in situ* in a diamond-anvil high-pressure camera at 21 GPa

TABLE 2. Pressure dependence of the unit cell volume of perovskite  $(Ca,Fe)(Si,Fe)O_3$  at room temperature

Pressure (GPa)	Unit cell volume (Å <sup>3</sup> )	
21	41.42	
17	42.39	
13	43.38	
10	44.01	

\* a = 3.460 (4)Å and V = 41.42 (14)Å<sup>3</sup>

The cubic cell parameter was determined as a function of pressure on decompression (Table 2). Comparison of the compressibility of the cubic  $Ca_3Fe_2Si_3O_{12}$  phase with that of  $CaSiO_3$  perovskite (Tamai and Yagi, 1989; Yagi *et al.*, 1989; see Fig. 3) shows that, at pressures >20 GPa, the unit cell is noticeably smaller than that of  $CaSiO_3$ 

perovskite, but with decreasing pressure its cell volume becomes gradually closer to that of CaSiO<sub>3</sub> perovskite. At ~11 GPa the cell volumes are almost equal, and below this pressure the cell volume of the new perovskite phase is greater than that of Ca-perovskite. Considering the ionic radii of Si<sup>4+</sup>, Fe<sup>3+</sup> (Fe<sup>3+</sup> > Si<sup>4+</sup>) and Ca<sup>2+</sup> (Fe<sup>3+</sup> < Ca<sup>2+</sup>), and taking account of charge balance, it is reasonable to suggest that this new high-pressure phase can be written as (Ca,Fe)(Si,Fe)O<sub>3</sub>, with ferric iron replacing Ca<sup>2+</sup> and Si<sup>4+</sup> in the CaSiO<sub>3</sub> perovskite structure (Fe<sup>3+</sup> + Fe<sup>3+</sup> = Si<sup>4+</sup> + Ca<sup>2+</sup>).



FIG. 3. Compressibilities of CaSiO<sub>3</sub> perovskite (after Tamai and Yagi, 1989) and (Ca,Fe)(Si,Fe)O<sub>3</sub> perovskite (this study: triangles).

The lattice parameter is calculated to be 3.460(4) Å at 21 GPa and room temperature. This yields a density of 5.09 g/cm<sup>3</sup>, which is ~13.6% greater than that of an isochemical andradite at the same pressure.

The diffraction pattern of (Ca,Fe)(Si,Fe)O<sub>3</sub> perovskite was followed upon release of pressure. Some runs were conducted at about 20 GPa and 30 GPa, and in all cases the recovered materials consisted of glass with a few weak lines corresponding to the andradite, which is present due to incomplete transformation. The zeropressure lattice parameter of the perovskite modification of (Ca,Fe)(Si,Fe)SiO<sub>3</sub> was not measurable experimentally. Extrapolation of lattice parameters from (Ca,Fe)(Si,Fe)SiO<sub>3</sub> obtained at different pressures gives a higher value than that of Ca-perovskite (CaSiO<sub>3</sub>) (Fig. 3). Because there was no pressure-transmitting medium present, pressure gradients must exist in the sample. Taking these gradients into account, and noting the difference in colour of andradite and the perovskite phase, the transformation pressure for andradite to perovskite is estimated at ~15-17 GPa. Based on previous data and our results on CaSiO<sub>3</sub>-Fe<sub>2</sub>O<sub>3</sub>, an increase in ferric iron content may increase this transition pressure under isothermal conditions (Fig.4). However, further studies on the CaSiO<sub>3</sub>-Fe<sub>2</sub>O<sub>3</sub> system will have to be carried out to clarify the position of the phase boundary and the stability field for Fe<sup>3+</sup>-bearing perovskite.

A series of studies on diamond inclusions derived from the deeper mantle have indicated that ferric iron can enter perovskite (Fe,Mg)SiO<sub>3</sub> and magnesiowüstite (Mg,Fe)O (McCammon et al., 1997; Harris et al., 1997; McCammon, 1997), and that the concentration of ferric iron is (positively) correlated to Al<sub>2</sub>O<sub>3</sub>-content. Our high-pressure experiments provide further very strong evidence that Fe<sup>3+</sup> can be incorporated into Ca-perovskite in the lower mantle. Previous electrical conductivity measurements on the lower mantle minerals magnesiowüstite and perovskite (Li and Jeanloz, 1990; Poirier and Peyronneau, 1992) provide conflicting results. The ambient temperature electrical conductivity of perovskite and a mixture of (Mg,Fe)SiO<sub>3</sub> perovskite and (Mg,Fe)O magnesiowüstite, measured by Li and Jeanloz (1990) is lower, by about three orders of magnitude, than that reported by Poirier and Peyronneau (1992). The results of Poirier and Peyronneau, extrapolated to lower mantle pressures, are in agreement with the



FIG. 4. Possible pressure-composition phase diagram for the garnet-perovskite transition (And: andradite).

value of 1 S/m at a depth of 1000 km, determined from the analysis of the transient and secular variations of geomagnetic field. In light of this result, it is of interest that Poirier and Peyronneau (1992) suggested a conduction mechanism in perovskite that involves electronic charge transfer between  $Fe^{2+}$  and  $Fe^{3+}$  to explain their high electrical conductivity. Based on the above data and conductivity mechanisms in semiconductors, it is both necessary and reasonable to deduce that another more highly conductive and interconnected phase exists in lower mantle (Li and Jeanloz, 1993). Thus, either Fe<sup>3+</sup> doping in mantle minerals or the presence of a new Fe<sup>3+</sup>-bearing perovskite phase could play an important role in the control of geophysical parameters, bearing in mind that the doping in semicoductors can induce huge increases in electrical conductivity. Although <10% Ca-perovskite can enter a lower mantle of pyrolite composition, Fe3+-rich Caperovskite might still provide an important control on geophysical evolution of lower mantle (just as, e.g. 4-5 wt.% Al<sub>2</sub>O<sub>3</sub> could: Wood and Rubie, 1996). There is no doubt that the presence of ferric iron can significantly affect the geochemical and geophysical properties of lower mantle, including its density and electrical conductivity.

In conclusion, we have demonstrated the existence of a perovskite type modification of  $Ca_3Fe_2Si_3O_{12}$  at high pressures, which supports the existence of ferric iron in the lower mantle, based on mantle xenoliths. As suggested elsewhere, it seems probable that  $(Ca,Fe)(Si,Fe)SiO_3$  perovskite is a significant constituent of the lower mantle, especially considering the electrical

conductivity of mantle mineral mixtures. In the light of these discoveries, it seems likely that the mineral species of the lower mantle possess more abundant ferric iron than ferrous iron.

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