Petrology and mineral chemistry of the Ascutney Mountain igneous complex

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ABSTRACT

The Ascutney Mountain igneous complex in southeastern Vermont is a Cretaceous member of the White Mountain plutonic-volcanic series. A subvolcanic complex, it consists of three stocks: gabbro-diorite, quartz syenite, and granite. A syenite porphyry ring dike rims a portion of the complex and contains large xenoliths of a unique breccia.

The gabbro-diorite consists of zoned plagioclase, augite, ferroan to magnesian hornblende, phlogopite, orthoclase, and quartz. A variety of syenites occur and consist of perthite (\pm albite \pm orthoclase), ferro-edenite, annite, quartz, Fe-rich augite, and fayalite. The granite consists of microperthite, orthoclase, albite, phlogopite, and edenite. Magnetite, ilmenite, apatite, titanite, and zircon are common accessory minerals.

The temperature of the syenite magma was 890–1000 °C, and f_{02} ranged from $10^{-13.7}$ to $10^{-12.9}$ bars, as determined from the assemblage quartz + magnetite + ilmenite + fayalite. Pressure is constrained to have been approximately 2 kbar. Biotite equilibria indicate that $f_{H_{20}}$ of the syenite magma was 3300–4800 bars ($a_{H_{20}} = 1.8-2.8$).

INTRODUCTION

The Ascutney Mountain igneous complex belongs to the White Mountain plutonic-volcanic series, a group of Mesozoic calc-alkalic to alkalic igneous complexes that occur in New Hampshire, southern Maine, and Vermont (Billings, 1928, 1943, 1956; Foland and Faul, 1977; McHone and Butler, 1984; Eby, 1987; Foland et al., 1985, 1988). Extensions of this series to the northwest and southeast, respectively, are the Monteregian Hills in the St. Lawrence Lowlands of Quebec and the New England Seamount Chain. (Eby, 1985; Gold, 1967; Philpotts, 1974; Fletcher et al., 1974; Bedard, 1988).

Ascutney Mountain is a classic locality for igneous petrology. It was as a result of his study of petrology and intrusive relations of the stocks there that Reginald Daly (1903) developed a theory of magmatic stoping as a mechanism for emplacement of large magma bodies. Subsequent work has focused on other mechanisms of intrusion (Chapman and Chapman, 1940), the Ascutney contact aureole (Nielson, 1973), the age of the rocks of the complex (Foland and Faul, 1977), and the isotopic composition of the constituent gabbro-diorite and granite (Foland et al., 1985, 1988). This paper reports on a field and petrologic study whose main purpose was to determine the physical conditions at the time of formation of the Ascutney complex. It is the first modern mineralogic and petrologic study of this classic field locality. Many of the igneous rocks and constituent minerals analyzed by electron microprobe in this study were collected and examined by R. A. Daly in his original investigation of Ascutney Mountain.

The Ascutney complex is located just west of the Connecticut River in southeastern Vermont (Fig. 1). Approximately 8 km (east-west) by 4 km (north-south) in size, the stocks and dikes of the complex were emplaced into Precambrian basement gneisses of the Chester dome and the overlying east-dipping Paleozoic metasediments. Superimposed on the regional metamorphic isograds is a well-developed contact metamorphic aureole around the main syenite stock (Nielson, 1973).

Radiometric ages obtained using the K-Ar method with biotite from intrusive rocks at Ascutney indicates that all the igneous rocks were emplaced in a short interval at 122 m.y. \pm 1.2 m.y. (Foland and Faul, 1977; Foland et al., 1985). Other White Mountain plutonic-volcanic series rocks were emplaced at three time periods, ~225 m.y., 180 m.y., and 120 m.y., with the majority occurring around 180 m.y. The origin of this series cannot be attributed to the westward movement of the North American plate over a single, fixed mantle hot spot because there is no linear correlation between geographic position and age for the White Mountain, Monteregian, and New England Seamount intrusions (Foland and Faul, 1977; Vink et al., 1985). Rather, the emplacement is attributed to intrusion along transform faults or fracture zones of a failed Mesozoic rift (Ballard and Uchupi, 1975).

ROCK TYPES AND PETROGRAPHY

The Ascutney Mountain complex consists of three stocks which are, in order of decreasing age, gabbro-diorite, syenite, and granite as well as a partial ring dike and numerous discordant dikes (Fig. 2).

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Gabbro-diorite

The gabbro-diorite stock along the westernmost extent of the intrusive complex forms the main mass of Little Ascutney Mountain. It consists of gabbro and diorite with varied amounts of augite, hornblende, and biotite.

Plagioclase, alkali feldspar, and quartz are the other major minerals of the mafic intrusion, in addition to augite, biotite, and hornblende. Zircon, titanite, apatite, magnetite, and ilmenite are accessory minerals. The rocks are coarse to medium grained. Hornblende occurs as rims around augite and also displays resorbed edges. Inclusions of almost all minerals are observed in euhedral biotite grains. The order of crystallization was: apatite, titanite, zircon, magnetite, ilmenite, plagioclase, pyroxene, hornblende, biotite, alkali feldspar, quartz. In the field, the presence of gabbro inclusions in diorite indicates the earlier emplacement of the gabbro.

Syenite

Syenite composes the main mass of Ascutney Mountain and forms a crescentic ring dike on the south side of Little Ascutney Mountain. Daly (1903) described the main stock as a composite of four rock types: hornblende-biotite nordmarkite (quartz-bearing alkali syenite), porphyritic hornblende-biotite-augite nordmarkite, alkaline granite, and monzonite.

The syenite is composed of porphyritic, seriate, or subhedral granular medium-grained aggregates of perthite, hornblende, and biotite. Minor constituents include quartz, augite, and fayalite. Apatite, zircon, titanite, monazite, ilmenite, and magnetite are accessory minerals. Unlike its occurrence in the gabbro-diorite, hornblende in the syenite appears to have crystallized later than biotite because biotite inclusions occur in euhedral hornblende crystals. The syenite dike of Little Ascutney Mountain is structurally and chemically different from the syenites of the main mountain; it is fine grained, porphyritic, and less Fe-rich than those on Ascutney Mountain proper.

A variety of xenoliths occurs within the syenite stock. Cognate xenoliths of gabbro and diorite are present in the syenite and are direct field evidence for the older age of emplacement of the mafic stock. Other small mafic inclusions abundant in the syenite consist of fine-grained plagioclase, hornblende, biotite, and quartz. Foland et al. (1985) suggest that these inclusions were stoped from overlying cogenetic volcanic rock as the magma was emplaced. Volcanic rock occurs as a screen in the western portion of the syenite stock and as smaller elliptical bodies north of the northern summit of Ascutney Mountain. Large country rock xenoliths also occur. Breccia xenoliths abundant in the syenite porphyry dike of Little Ascutney Mountain were the subject of a detailed study by the author (Schneiderman, 1989). It is the presence of such xenoliths in the syenite at Ascutney, particularly at Crystal Cascade, that led Daly (1903) to formulate his theory of magmatic stoping as a mechanism for the emplacement of intrusive bodies.



Fig. 1. Generalized geologic map of southeastern Vermont and southwestern New Hampshire showing location of the Ascutney igneous complex. Modified from Billings (1956), Doll et al. (1961), and Downie (1982).

Granite

Granite intrudes the southeast portion of the main syenite stock at Ascutney. It is medium to coarse grained and subhedral granular to subporphyritic. Microperthite, orthoclase, albite, quartz, and biotite are the major constituents. Hornblende is only rarely observed. Apatite, zircon, titanite, magnetite, and ilmenite are accessory minerals. Rare hornblende displays resorbed edges and biotite occurs as large, euhedral grains. This biotite granite is comparable to the Conway granite of the White Mountain plutonic-volcanic series in New Hampshire.

Dike rocks

Daly (1903) mapped four types of dike rocks at Ascutney Mountain: windsorite, paisanite, muscovite aplite, and lamprophyres. Balk and Krieger (1936) recognized devitrified felsites containing spherulites.

Daly (1903) defined the rock type windsorite to describe a leucocratic variety of quartz monzonite found near Windsor, Vermont. The rock type is characterized



Fig. 2. Generalized geologic map of the Ascutney igneous complex. Modified from Daly (1903), Chapman and Chapman (1940), and Nielson (1973).

by high alkali content (K > Na), low Ca, Fe, and Mg (Daly, 1903). Two such dikes cut the eastern portion of the gabbro-diorite stock. They consist of plagioclase (oligoclase to andesine), orthoclase, microperthite, and bio-tite. Some of the plagioclase is rimmed by alkali feldspar. Quartz occurs interstitially and augite and hornblende are present though rare.

Fine-grained porphyritic dikes consisting of microperthite, orthoclase, quartz, and alkali-rich hornblende (paisanite) cut the gabbro-diorite, syenite ring dike, and the main syenite stock. Petrographically they resemble fine-grained phases of the main syenite stock. Like the syenite stock they contain abundant glomeroporphyritic aggregates of hornblende and biotite.

Muscovite aplite cuts the main syenite stock. It consists of quartz, orthoclase, albite, microperthite, and muscovite. Miarolitic cavities are common and contain terminated quartz crystals and books of muscovite.

Two other types of dikes, camptonite and diabase, cut the gabbro-diorite, syenite stock, and the country rock. The camptonite dikes are dark colored and fine grained and consist of plagioclase and hornblende that is altered to chlorite, epidote, calcite, and quartz. The diabase dikes, like those observed throughout the Connecticut River Valley, consist primarily of plagioclase and interstitial augite. Pyrite and magnetite are the main accessory minerals.

Aphanitic, felsite dikes studied by Balk and Krieger (1936) cut country rock gneisses, gabbro-diorite, and

TABLE 1. Representative pyroxene analyses

	Gabbro	-diorite	Sy	renite
	5237-7	21361-5	2187-7	AS84113-4
SiO	52.88	52.33	50.61	47.66
TiO	0.26	0.21	0.41	1.71
Al ₂ O ₃	0.88	0.93	2.78	5.46
Cr.O.	0.00	0.01	0.02	0.04
MaO	14.17	13.89	9.91	12.03
FeO	9.68	9.70	20.56	11.15
MnO	0.61	0.47	0.65	0.38
CaO	21.10	21.44	11.44	19.16
Na ₂ O	0.42	0.36	1.08	1.01
Total	100.00	99.34	97.46	98.60
	Norm	alized to four o	ations	
Si	1.97	1.97	2.00	1.81
Ti	0.01	0.01	0.01	0.05
AI	0.04	0.04	0.13	0.24
Cr	0.00	0.00	0.00	0.00
Ma	0.79	0.78	0.59	0.68
Fe	0.30	0.31	0.68	0.35
Mn	0.02	0.02	0.02	0.01
Ca	0.84	0.86	0.49	0.78
Na	0.03	0.03	0.08	0.07

TABLE 2. Representative olivine analyses

	Sye	nite
	5234-8	5234-9
SiO2	30.87	30.85
TiO	0.02	0.03
Al ₂ O ₃	0.00	0.00
Cr ₂ O ₃	0.03	0.04
MgO	1.78	1.76
FeO	64.28	63.58
MnO	4.11	4.42
NiO	0.03	0.03
CaO	0.10	0.12
Total	101.22	100.83
1	Normalized to three cation	15
Si	1.02	1.02
Ti	0.00	0.00
AI	0.00	0.00
Cr	0.00	0.00
Mg	0.09	0.09
Fe	1.77	1.76
Mn	0.12	0.12
Ni	0.00	0.00
Ca	0.00	0.00

quartz syenite of Little Ascutney. They are distinguished by flow-banding, spherulites, and a devitrified groundmass. The spherulites occur in clusters that are roughly parallel to the flow-banding and are composed of microlites arranged in a radial fashion.

MINERAL COMPOSITIONS

Mineral compositions were determined with an automated three spectrometer Cameca MBX electron microprobe at Harvard University. The data reported in this study represent complete microprobe analyses for up to 14 elements: Si, Al, Ti, Fe, Mg, Mn, Zn, Cr, Ca, Ba, K, Na, F, Cl. The spectrometers were prioritized to measure Na, K, F, and Cl first during the analysis of biotite and amphibole. Oxide weight percents were converted from X-ray intensities using the procedures of Bence and Albee (1968) and the correction factors modified from Albee and Ray (1970). Operating conditions were 15 kV accelerating potential and 15 nA beam current. A set of simple oxides and silicates was used for primary standards. Counting statistics were set for 1% standard deviation.

For exsolved feldspars, bulk compositions were obtained for cryptoperthite with regular $1-2 \mu m$ exsolution lamellae by systematic analysis with a 16 $\mu m \times 16 \mu m$ raster beam over 20–25 microprobe areas per grain. Compositions of lamellae and host were determined for microperthite (5–7 μm lamellae) grains using a 1–2 μm point beam.

Note that, in the tables of representative mineral analyses, the amalgam of numbers and letters that precedes a hyphen, if present, is the rock sample number; the number after the hyphen is the number of the analysis. Therefore, for example, the minerals for which analyses were obtained, fayalite analysis 5234-8, plagioclase 5234-5, alkali feldspar 5234-5, and Fe-Ti oxides 5234-7, all come from the same rock sample.

Pyroxene

Clinopyroxene is abundant in the gabbro-diorite of Little Ascutney Mountain and sometimes constitutes as much as 20% of the rock. It is a minor mineral phase in the syenite, subordinate to biotite and hornblende, and is not found in the granite. It is colorless to pale greenish brown, lacks pleochroism, and is simply twinned. Representative electron microprobe analyses are given in Table 1.

TABLE 3. Representative plagioclase analyses

		Gabbro	o-diorite						Granite				
	Core	Rim	Core	Rim	Syenite			Core	Rim	Core	Rim		
	21361	21361	3480	3480	5234-5	AS841A-8	AS84113-1	21369	21369	5239	5239		
SiO ₂	48.20	57.61	47.83	59.08	61.13	65.34	66.44	62.34	66.37	65.90	66.22		
Al ₂ O ₃	32.63	26.40	32.98	25.53	24.53	21.19	20.96	22.91	20.61	20.98	20.79		
TiO ₂	0.09	0.06	0.04	0.01	0.05	0.00	0.03	0.05	0.01	0.04	0.03		
FeO	0.35	0.24	0.39	0.23	0.14	0.27	0.14	0.20	0.21	0.18	0.16		
MgO	0.01	0.00	0.01	0.01	0.00	0.01	0.00	0.02	0.02	0.01	0.00		
BaO	0.02	0.04	0.01	0.01	0.13	0.03	0.00	0.09	0.00	0.01	0.01		
SrO	0.16	0.15	0.20	0.17	0.11	0.07	0.01	0.19	0.03	0.09	0.06		
CaO	15.78	8.30	16.04	7.01	5.72	2.21	1.66	4.32	1.51	1.88	1.54		
Na₂O	2.47	6.56	2.27	7.15	8.26	9.46	10.66	8.46	10.35	10.23	10.38		
K ₂ O	0.08	0.31	0.08	0.55	0.36	1.37	0.40	0.67	0.48	0.58	0.42		
Total	99.79	99.67	99.85	99.70	100.43	99.95	100.30	99.35	99.59	99.90	99.61		
			Atom	nic proportio	ns calculate	d on the bas	is of eight O a	toms					
Si	2.22	2.59	2.20	2.65	2.71	2.89	2.91	2.79	2.93	2.91	2.92		
AI	1.77	1.40	1.79	1.35	1.28	1.11	1.08	1.21	1.07	1.09	1.08		
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00		
Fe	0.01	0.01	0.02	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01		
Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00		
Ba	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00		
Sr	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00		
Ca	0.78	0.40	0.79	0.34	0.27	0.11	0.08	0.21	0.07	0.09	0.07		
Na	0.22	0.57	0.20	0.62	0.71	0.81	0.91	0.73	0.89	0.87	0.89		
ĸ	0.01	0.02	0.01	0.03	0.02	0.08	0.02	0.04	0.03	0.03	0.02		
Σ cation	5.01	5.00	5.01	5.00	5.00	5.01	5.01	4.99	4.99	5.00	5.00		



Fig. 3. Ternary projection of feldspar compositions from gabbro-diorite, syenite, and granite from Ascutney, Vermont. (A) Compositions of coexisting unzoned plagioclase and unzoned alkali-feldspars. Triangle symbols represent coexisting feldspars from granite. Other symbols represent coexisting feldspars form syenite. (B) Core and rim compositions for zoned plagioclase and alkali feldspar grains and compositions of coexisting plagioclase rims and alkali feldspar rims.

Fayalite

Fayalite occurs as an accessory mineral in the syenite. Representative electron microprobe analyses (Table 2) show that its composition is $Fa_{89}Fo_5Te_6$. It is commonly rimmed by amphibole and biotite and is usually altered to iddingsite.

Feldspars

Both plagioclase and alkali feldspar occur in all the intrusive rocks of the Ascutney complex. Plagioclase is the predominant feldspar in the gabbro-diorite. The syenite is hypersolvus although discrete grains of potassium feldspar and albite are observed occasionally. In the granite, plagioclase and potassic feldspar occur in subequal proportions. The feldspars occur as subhedral laths and anhedral grains ranging in size from 0.5 mm to 4 mm. Plagioclase is twinned according to albite and Carlsbad twin laws. Myrmekitic intergrowths of wormy quartz in plagioclase are observed between plagioclase and potassium feldspar in the gabbro-diorite.

In the gabbro-diorite and granite, plagioclase is normally zoned with cores of $X_{An} = 0.40-0.50$ and $X_{An} = 0.10-0.15$ and rims of $X_{An} = 0.20$ and $X_{An} = 0.0-0.50$, respectively (Table 3). In the syenite, small, twinned grains of albite ($X_{An} = 0.0-0.10$) are mantled by potassic feldspar. K content of plagioclase cores is highest in syenite (6 wt%). Unzoned plagioclase in the syenite is $X_{An} = 0.08$ and that in the granite is $X_{An} = 0.10$ (Fig. 3a).

Alkali feldspar is present as crypto- and microperthite, orthoclase, and microcline. Both the perthite and orthoclase have Carlsbad twinning and the microcline displays characteristic grid twinning. Small unzoned alkali feldspar grains range in composition from $X_{or} = 0.40$ to $X_{or} = 0.79$ (Fig. 3a). Large alkali feldspars are normally zoned from sodic cores to potassic rims (Fig. 3b). Cores of many perthite grains are weathered and altered to fine-grained sericite. Granophyric and micrographic intergrowths of quartz in perthite are common. Representative electron microprobe analyses and structural formulae are given in Table 4.

Amphibole

Amphibole in all intrusive rocks of the Ascutney complex is strongly pleochroic from light greenish brown to gray olive-green, displays well developed cleavage, and is frequently twinned parallel to (100). It frequently rims augite and contains inclusions of ilmenite, magnetite, zircon, and apatite.

Amphiboles from Ascutney intrusions are calcic: ^{M4}(Ca + Na) \geq 1.34; ^{M4}Na < 0.67 (Leake, 1978). In the syenite, amphibole is ferro-edenite and ferro-edenitic hornblende; the granite contains edenite; in contrast, the gabbro-diorite and the porphyritic syenite at Little Ascutney have ferro-hornblende and magnesio-hornblende (Table 5). Compositional variations within amphibole are as follows: wt% Al₂O₃ correlates inversely with SiO₂ whereas it correlates positively with Na₂O, K₂O, and TiO₂, indicating coupled substitutions.

Figure 4 shows cationic Mg/(Mg + Fe^{2+} + Mn) vs. ¹⁶Al/¹⁴Al of the Ascutney amphiboles. Amphiboles in the gabbro-diorite and granite are relatively enriched in Mg compared to other amphiboles in the complex. An inverse correlation between Mg and ¹⁴Al of amphibole in syenite indicates a tschermak substitution.

Bedard (1988) showed that Monteregian Hills and White Mountain series amphiboles plot in distinct fields with regard to Ca + $^{[4]}Al$ and Si + Na + K. Ascutney amphiboles fall within Bedard's field of White Mountain series amphiboles indicating that the gabbro-diorite, sy-enite, and granite crystallized from silica-saturated magmas.

						Syenite						
	Gabbro-	-			46.0	24.10		AS	348-13		Granite	
	3480-7	3501-10	5234-5	AS841A-8	Core	Rim	Host	Host rim	Lamellae core	Lamellae rim	21369	5239-7
SiO	64.56	65.93	65.47	65.55	65.28	65.36	65 17	65 34	67.47	67 58	65.83	65.09
ALO	18.34	18 95	19 43	18 76	18 72	18.67	18.57	18 53	10.08	10.99	19.93	18.46
TiO	0.02	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.06	0.05
FeO	0.12	0.19	0.27	0.22	0.06	0.06	0.05	0.00	0.07	0.00	0.16	0.16
MaO	0.00	0.03	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01
BaO	0.40	0.05	0.07	0.13	0.07	0.05	0.04	0.07	0.00	0.00	0.12	0.17
SrO	0.16	0.01	0.03	0.06	0.02	0.03	0.04	0.06	0.02	0.03	0.02	0.08
CaO	0.06	0.20	0.71	0.26	0.13	0.08	0.05	0.02	0.53	0.50	0.25	0.12
Na ₂ O	1.52	4.52	4.76	3.26	3.03	2.37	1.32	1 40	10.39	9.95	5 52	2 92
K ₂ Ô	13.61	10.08	9.11	11.50	12 16	13 15	14.58	14 55	1.50	2 21	8.07	11.57
Total	98.79	99.97	99.86	99.76	99.47	99.77	99.82	100.02	99.96	100 21	98.86	98 63
			A	tomic propor	tions cald	culated on	the basis o	f eight O at	toms	100.21	00.00	00.00
Si	3.00	2.99	2.96	2.99	2.99	3.00	3.00	3.00	2.97	2.97	2.99	3.00
AI	1.01	1.01	1.04	1.01	1.01	1.01	1.01	1.00	1.04	1.03	1.01	1.00
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Fe	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01
Mg	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Ba	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Ca	0.00	0.01	0.03	0.01	0.01	0.00	0.00	0.00	0.03	0.02	0.01	0.01
Na	0.14	0.40	0.42	0.29	0.27	0.21	0.12	0.13	0.89	0.85	0.49	0.26
K	0.81	0.58	0.53	0.67	0.71	0.77	0.86	0.85	0.08	0.12	0.47	0.68
Σ cation	4.97	5.00	4.99	4.98	4.99	4.99	4.99	4.99	5.00	5.00	4.98	4.97

TABLE 4. Representative alkali feldspar analyses

Biotite

Biotite occurs in all intrusive rocks of the Ascutney complex in glomeroporphyritic aggregates with amphibole, augite, ilmenite, magnetite, zircon, and apatite. It is strongly pleochroic. Also, well-developed pleochroic haloes surround zircon inclusions. In the gabbro-diorite and granite biotite is subhedral, whereas in the syenite, it forms irregularly shaped grains and occurs as inclusions in hornblende, indicating its early crystallization.

Biotite structural formulae were calculated on the basis of 11 O atoms since high valence cations substitute for divalent ones in Ascutney biotite and result in octahedral vacancies (Table 6).

Biotite from gabbro-diorite and granite has $Fe/(Fe^{2+} + Mg)$ between 0.41 and 0.45, biotite from the synite ring dike ranges from 0.50 to 0.72, and biotite from the synite stock is most Fe rich, $Fe/(Fe^{2+} + Mg) = 0.62$ to 0.85. Biotite from all rock types has low Al but high K and F contents.

The biotite compositions reflect both tschermak and octahedral vacancy substitutions. The positive correlation between ^[4]Al and Ti indicates a component of titanium tschermak substitution. Also, an apparent correlation between higher Ti contents and lower octahedral site occupancies suggests the coupling of Ti substitutions with an octahedral vacancy $(2R^{2+} = R^{4+} + \Box)$.

Fe-Ti oxides

Magnetite and ilmenite occur in all intrusive rocks of the complex. They occur in subequal proportions of the gabbro-diorite and granite, but ilmenite is more abundant than magnetite in the syenite. Both minerals occur as discrete, subhedral to anhedral grains distributed equally throughout the rock, in glomeroporphyritic aggregates, or as inclusions in augite, hornblende, and biotite. Magnetite flakes form rims around biotite adjacent to alkali feldspar, a texture that may result from the reaction annite + ilmenite = potassium feldspar + ulvospinel + H_2O (Diamond and Frost, 1988). Table 7 shows representative microprobe analyses of coexisting oxides.

GEOBAROMETRY AND GEOTHERMOMETRY

Pressure

Andalusite in the Ascutney contact aureole indicates that pressure during emplacement of the intrusive complex was no greater than 3.8 kbar (Holdaway, 1971; Nielson, 1973; Schneiderman, 1987). Compositions and textures of the mafic silicates in the syenite supply lower pressure estimates.

Annite and quartz rim fayalite which is adjacent to potassium feldspar (Fig. 5). Assuming that the melting curve of the Ascutney quartz syenite is comparable to the granite minimum melting curve, the intersection of this curve with the stability curve for the reaction

$$\begin{array}{rcl} 2KFe_{3}AlSi_{3}O_{10}(OH)_{2} + 3SiO_{2} &=& 2KAlSi_{3}O_{8}\\ & & \\ annite & & \\ quartz & & \\ potassium feldspar \\ & +& 3Fe_{2}SiO_{4} + 2H_{2}O\\ & & \\ & & \\ fayalite & & \\ vapor \end{array}$$

indicates a pressure of at least 1.5 kbar at the time of intrusion (Fig. 6). Although Ca and Na in the A-site and octahedral Al decrease the stability of annite (Wones and Eugster, 1965) they are not substantial components of the Ascutney biotite. However, Mg, Ti, and F, all present in Ascutney biotite, expand the P-T field of Fe-rich biotite (Rutherford, 1969; Eugster and Wones, 1962).

TABLE 5. NE	epresentative	ampriluoi	e analyses								
	G	abbro-diori	te			Syenite	Granite				
	21361-1	3480-2	5237-3	AS848-4	AS8444-7	3490-2	2187-4	5250-4A	21369-1	21369-2A	21369-2E
SiO ₂	41.87	47.14	48.01	43.14	46.64	46.44	41.80	42.56	48.03	48.04	47.77
Al ₂ O ₃	10.90	6.12	6.15	7.57	5.44	5.35	8.00	7.13	5.40	5.46	5.38
TiO ₂	4.35	1.23	0.99	0.29	0.20	0.86	1.60	1.58	0.81	1.15	1.05
FeO*	13.01	14.18	13.83	27.99	23.43	22.54	29.03	27.18	14.47	38.88	14.46
MgO	12.57	13.72	14.19	4.22	7.59	8.22	3.66	4.88	13.67	13.85	13.55
MnO	0.27	0.34	0.49	1.27	0.99	1.08	0.98	0.94	1.18	1.05	1.17
ZnO	0.01	0.01	0.04	0.05	0.06	0.06	0.06	0.08	0.04	0.02	0.04
Cr ₂ O ₃	0.02	0.01	0.00	0.00	0.01	0.00	0.02	0.01	0.01	0.03	0.01
CaO	11.31	11.71	10.77	10.33	11.68	10.56	10.27	9.90	10.99	10.96	11.05
BaO	0.14	0.03	0.03	0.02	0.02	0.01	0.02	0.03	0.03	0.02	0.00
Na ₂ O	2.35	1.03	1.28	1.91	1.28	1.58	2.21	2.18	2.17	2.26	2.23
K,Ō	1.04	0.67	0.53	0.97	0.70	0.71	1.24	1.09	0.79	0.72	0.74
F	0.20	0.22	0.22	0.66	0.64	0.70	0.92	1.15	1.62	1.64	1.52
CI	0.03	0.15	0.11	0.42	0.18	0.20	0.53	0.35	0.07	0.10	0.10
O = F + CI	-0.09	-0.13	-0.12	-0.37	-0.31	-0.34	0.51	-0.56	-0.70	-0.71	-0.66
Total	97.98	96.43	96.52	98.49	98.54	97.97	99.83	98.50	98.58	98.47	98.41
		Atomic p	proportions	calculated o	on the basis	of 13 catio	ons excludi	ng Ca, Na, E	Ba, K		
Si	6.20	6.97	7.00	6.75	7.13	7.07	6.54	6.66			
[4] A]	1.80	1.03	1.00	1.25	0.98	0.94	1.46	1.32	7.05	7.06	7.04
Σ T-site	8.00	8.00	8.00	8.00	8.01	8.01	8.00	7.98	0.94	0.94	0.93
[6]A]	0.10	0.04	0.06	0.14	0.11	0.02	0.02	0.00	0.00	0.01	0.00
Ti	0.48	0.14	0.11	0.03	0.02	0.10	0.19	0.19	0.09	0.13	0.12
Fe ^{3+**}	0.27	0.58	0.90	0.81	0.36	0.67	0.70	0.78	0.55	0.46	0.49
Cr ³⁺	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mg	2.77	3.03	3.08	0.98	1.73	1.86	0.85	1.14	2.99	3.03	2.98
Fe ²⁺	1.34	1.18	0.79	2.86	2.64	2.20	3.10	2.78	1.23	1.25	1.29
Zn	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00
Mn	0.03	0.04	0.06	0.17	0.13	0.14	0.13	0.10	0.14	0.12	0.12

5.00 0.00

1.91

0.09

2.00

0.29

0.14

0.00

0.43

60.41

TABLE 5. Representative amphibole analyses

Compositions of amphibole in the syenite approach those of hastingsite. Near FMQ conditions, the stability curve for the reaction hastingsite = magnetite + hedenbergite_{ss} + plagioclase + garnet_{ss} + fluid intersects the granite minimum melting curve at ~2.3 kbar (Thomas, 1977, 1982). Substitutions of Mg, F, and K shift the stability curve to higher temperatures toward that of annite, indicating pressures less than 2.3 kbar (Cameron and Gibbs, 1973; Helz, 1979).

5.01

0.00

1.86

0.14

2.00

0.15

0.13

0.00

0.28

28.03

5.00

0.00

1.68

0.32

2.00

0.05

0.10

0.00

0.15

20.41

5.00

0.00

1.73

0.27

2.00

0.31

0.19

0.00

0.50

74.48

4.99

0.00

1.79

0.21

2.00

0 46

0.20

0.01

0.67

32.60

** Fe³⁺ calculated by charge balance for 23 O atoms.

* Total Fe measured as FeO.

Biotite inclusions in hornblende indicate that the thermal stability of hornblende did not exceed that of the annite. In contrast, gabbro-diorite and granite contain biotite with hornblende inclusions. The inclusions are magnesiohastingsite which has a greater thermal stability than hastingsite (Semet and Ernst, 1981; Thomas, 1982).

Temperature

Temperatures ranging from 550 to 725 °C were determined for the syenite samples using the Stormer (1975) and Powell and Powell (1977) geothermometers. However, it seems likely that these temperatures only reflect temperatures of subsolidus recrystallization and reequilibration.

One syenite sample (5234) contained the assemblage quartz + fayalite + ilmenite + magnetite and thus it was possible to apply the QUIIF equilibrium technique of Frost et al. (1988) to estimate temperature of the syenite. Assuming 2 kbar pressure and $X_{ilm} = 0.97$, the calculated temperature of the syenite is 890–1000 °C. This temperature range is corroborated by the contact aureole assemblage cordierite + spinel + corundum + sillimanite. Also, fayalite-bearing rhyolites from other quartz-saturated igneous rocks show similarly high temperatures (Frost et al., 1988).

Oxygen fugacity

5.00

0.00

1.72

0.28

2.00

0.18

0.14

0.00

0.32

54.19

5.00

0.00

1.72

0.28

2.00

0.39

0.25

0.00

0.64

78.48

5.00

0.03

1.66

0.31

2.00

0.35

0.22

0.00

0.57

70.92

5.00

0.01

173

0.26

2.00

0.36

0.15

0.00

0.51

29.15

5.00

0.01

1.73

0.26

2.00

0.38

0.14

0.00

0.52

29.21

5.00

0.03

1.75

0.22

2.00

0.42

0.14

0.00

0.56

30.21

Coexisting oxides in the syenite occur as discrete grains. However, like the feldspars, the oxides have undergone reequilibration. Therefore, f_{O_2} could not be determined using the Fe-Ti oxide oxybarometer (Buddington and Lindsley, 1964; Spencer and Lindsley, 1981). It was possible, though, to use the QUIIF assemblage of syenite

Σ C-site

Σ B-site

Σ A-site

100 × XFe2+

Mn

Ca

Na

Na

Ba

κ



Fig. 4. $Mg/(Mg + Fe^{2+} + Mn)$ vs. ⁽⁶⁾Al/⁽⁴⁾Al for amphiboles in Ascutney intrusions.

sample 5234 to estimate f_{O_2} (Frost et al., 1988). Once again, taking $X_{ilm} = 0.97$ and 2 kbar pressure, the f_{O_2} ranges from $10^{-13.7}$ bars to $10^{-12.9}$ bars, just below the FMQ buffer.

H₂O activity

The temperature and f_{02} estimates were used with the compositions of alkali feldspar, magnetite, and biotite,

assuming ideal solution, to calculate f_{H_2O} . Modifying the equation of Wones (1981) to include the effects of F substitution for OH on the activity of annite (Czmanske and Wones, 1973) and using it with the estimated activities of annite in biotite (0.85), KAlSi₃O₈ in alkali feldspar (0.6), Fe₃O₄ in spinel (0.86), T = 890-1000 °C, and $f_{O_2} =$ 10^{-13,7}-10^{-12.9} bars, syenite sample 5234 gives values of $f_{\rm H_{2O}}$ ranging from approximately 3300-4800 bars. [The crystal chemical models of Waldbaum and Thompson (1969) and Stormer (1983) were used to calculate the activities of KAlSi₃O₈ in feldspar and Fe₃O₄ in spinel, respectively.] This $f_{\rm H_{2}O}$ is considered to be a maximum because the biotite from Ascutney syenite most likely contains Fe³⁺ and the activity of annite in biotite would be less than 0.85. Using the tabulations of Burnham et al. (1969) and taking the pressure at the time of emplacement to be approximately 2 kbar as explained previously, the calculated H₂O activity ranges from 1.8–2.8. The range of values reflects the temperature range of 890-1000 °C. This high H₂O activity indicates that the magma was fluid saturated.

This is not a surprising result. Early crystallization of anhydrous minerals would contribute to increased $a_{H_{2}O}$

TABLE 6. Representative biotite analyses

Ga	abbro-dior		Syenite								Granite		
3480 5 5237 4 21261 1		AC040 2	2400 5	01067 1	AS	0107.1	5250.1	5024 4	5220 4	21369-	AC040A		
5400-5	5257-4	21301-1	A3040-2	3490-5	21307-1	04 I A-4	2107-1	5250-1	5234-4	5239-4	ЗА	A3042A	
36.09	36.11	36.40	33.79	34.63	34.87	35.44	34.89	34.97	35.01	38.83	38.23	38.93	
13.59	13.55	13.47	12.81	13.14	12.92	12.47	11.97	11.75	11.83	11.46	12.17	11.83	
4.66	4.94	4.88	3.71	4.21	6.07	4.21	3.32	2.52	4.02	1.92	3.17	2.26	
18.03	17.54	18.22	31.03	26.57	24.72	28.73	33.22	32.95	32.41	17.55	16.21	15.05	
12.82	13.02	12.47	3.83	7.21	7.68	5.81	3.34	3.88	3.31	14.28	14.62	16.00	
0.13	0.17	0.15	0.70	0.47	0.37	0.35	0.46	0.34	0.52	1.14	0.73	0.86	
0.01	0.02	0.03	0.11	0.08	0.07	0.18	0.08	0.07	0.16	0.09	0.03	0.06	
0.02	0.02	0.03	0.11	0.15	0.05	0.03	0.01	0.06	0.02	0.03	0.01	0.04	
0.00	0.00	0.00	0.04	0.02	0.03	0.04	0.02	0.04	0.01	0.00	0.00	0.01	
0.72	0.59	0.37	0.08	1.71	1.90	0.10	0.06	0.10	0.11	0.03	0.12	0.07	
0.10	0.17	0.15	0.07	0.09	0.44	0.37	0.07	0.06	0.08	0.16	0.22	0.24	
9.53	9.60	9.60	9.11	8.85	8.46	8.68	9.19	9.20	9.28	9.86	9.79	10.04	
0.39	0.35	0.31	0.70	0.94	0.96	1.69	1.04	0.96	1.04	3.04	2.61	3.56	
0.27	0.23	0.14	0.49	0.36	0.15	0.58	0.83	1.10	0.57	0.05	0.09	0.13	
96.36	96.31	96.22	96.58	98.43	98.69	98.68	98.50	98.00	98.37	98.44	98.00	99.08	
-0.23	-0.20	-0.16	-0.41	-0.48	-0.44	-0.84	-0.63	-0.65	-0.57	-1.29	-1.12	-1.53	
3.60	3.63	3.66	3.15	3.18	3.24	2.75	2.92	2.88	2.99	2.43	2.63	2.21	
99.73	99.74	99.72	99.32	101.13	101.49	100.59	100.79	100.23	100.79	99.58	99.51	99.76	
		A	tomic prop	ortions c	alculated	on the bas	sis of 11	O atoms					
2.76	2.75	2.77	2.77	2.74	2.72	2.82	2.84	2.87	2.84	2.96	2.90	2.93	
1.22	1.22	1.21	1.23	1.23	1.19	1.17	1.15	1.13	1.13	1.03	1.09	1.05	
0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	
0.23	0.28	0.28	0.23	0.25	0.36	0.25	0.20	0.16	0.25	0.11	0.18	0.13	
1.15	1.12	1.16	2.13	1.76	1.61	1.91	2.26	2.26	2.20	1.12	1.03	0.95	
0.01	0.01	0.01	0.05	0.03	0.02	0.02	0.03	0.02	0.04	0.07	0.05	0.06	
1.46	1.48	1.42	0.47	0.85	0.89	0.69	0.41	0.47	0.40	1.62	1.65	1.80	
0.00	0.00	0.00	0.01	0.01	0.00	0.01	0.01	0.00	0.01	0.01	0.00	0.00	
0.00	0.00	0.00	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
0.02	0.02	0.01	0.00	0.05	0.06	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
0.02	0.03	0.02	0.01	0.01	0.07	0.06	0.01	0.01	0.01	0.02	0.03	0.04	
0.93	0.93	0.93	0.95	0.89	0.84	0.88	0.95	0.96	0.96	0.96	0.95	0.96	
0.09	0.08	0.08	0.18	0.24	0.24	0.43	0.27	0.25	0.27	0.73	0.63	0.85	
0.04	0.03	0.02	0.07	0.05	0.02	0.08	0.12	0.15	0.08	0.01	0.01	0.02	
2.85	2.89	2.87	2.91	2.91	2.88	2.88	2.91	2.92	2.90	2.93	2.94	2.94	
0.44	0.43	0.45	0.82	0.67	0.64	0.74	0.85	0.83	0.85	0.41	0.38	0.34	
	36.09 13.59 4.66 18.03 12.82 0.13 0.01 0.02 0.00 0.72 0.10 9.53 0.39 0.27 96.36 -0.23 3.60 99.73 2.76 1.22 0.00 0.23 1.15 0.01 1.46 0.00 0.23 1.15 0.01 1.46 0.00 0.23 1.15 0.01 0.02 0.02 0.02 0.03 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.02 0.00 0.27 96.36 0.02 0.02 0.00 0.23 0.01 0.02 0.00 0.02 0.02 0.02 0.02 0.00 0.27 96.36 0.02 0.02 0.00 0.23 0.01 0.02 0.02 0.02 0.02 0.02 0.02 0.02	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{ccccccccccccccccccccccccccccccccccc$					

	Gabbro-diorite				Sye	enite				Gran	ite	
	5237 mag	5237 ilm	5234-7 mag	5234-7 ilm	5240 mag	5240 ilm	5250 mag	5250 ilm	AS-84-1A mag	AS-84-1A ilm	5239 mag	5239 ilm
SiO ₂	0.08	0.02	0.10	0.03	0.10	0.07	0.10	0.09	0.13	0.06	0.06	0.06
TiO,	0.31	46.30	4.43	49.01	0.62	48.44	5.10	48.97	5.51	48.98	0.26	56.18
Al ₂ O ₃	0.46	0.03	0.47	0.02	0.16	0.04	0.69	0.08	0.68	0.04	0.08	0.04
V2O3	0.50	0.43	0.10	0.33	0.20	0.32	0.11	0.32	0.08	0.25	0.12	0.40
Cr ₂ O ₃	0.13	0.05	0.04	0.02	0.08	0.03	0.07	0.00	0.00	0.01	0.04	0.04
Fe ₂ O ₃	68.12	13.12	60.14	3.37	67.63	4.21	58.81	3.63	58.08	4.53	68.58	0.00
FeO	31.76	38.83	35.27	43.55	31.54	42.30	35.87	43.56	36.15	42.60	31.53	37.47
MnO	0.03	1.58	0.28	2.45	0.11	3.11	0.32	2.28	0.41	2.67	0.03	0.27
MgO	0.07	0.79	0.00	0.03	0.03	0.08	0.00	0.01	0.00	0.00	0.00	0.00
CaO	0.01	0.02	0.02	0.05	0.21	0.04	0.03	0.01	0.03	0.02	0.00	0.04
ZnO	0.12	0.00	0.05	0.02	0.03	0.05	0.20	0.08	0.29	0.16	0.02	0.04
NiO	0.00	0.19	0.04	1.88	0.01	1.89	0.01	1.65	0.01	1.25	0.03	1.94
Total	101.58	101.36	100.93	100.75	100.72	100.57	101.31	100.68	101.36	100.57	100.75	96.48
			Magnetite	e normalize	ed to 3 cati	ons; ilmen	ite normali:	zed to 2 ca	ations			
Si	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00
Ti	0.01	0.87	0.13	0.93	0.02	0.92	0.15	0.93	0.16	0.93	0.01	0.12
Al	0.02	0.00	0.02	0.00	0.01	0.00	0.03	0.00	0.03	0.00	0.00	0.00
V	0.02	0.01	0.00	0.01	0.01	0.01	0.00	0.01	0.00	0.01	0.00	0.01
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Fe ³⁺	1.94	0.25	1.71	0.06	1.94	0.08	1.67	0.07	1.65	0.09	1.97	0.00
Fe ²⁺	1.00	0.81	1.12	0.92	1.01	0.89	1.13	0.92	1.14	0.90	1.01	0.83
Mn	0.00	0.03	0.01	0.05	0.00	0.07	0.01	0.05	0.01	0.06	0.00	0.01
Mg	0.00	0.03	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Ca	0.00	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Zn	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.01	0.00	0.00	0.00
Ni	0.00	0.00	0.00	0.02	0.00	0.02	0.00	0.02	0.00	0.01	0.00	0.02
X _{usp} ; X _{ilm}	0.01	0.87	0.13	0.97	0.02	0.96	0.15	0.96	0.16	0.95	0.01	1.00
Note: Tota	I Fe measure	ed as FeO; I	Fe ³⁺ and mo	le fractions	recalculate	d according	g to Storme	er (1983).				

TABLE 7. Representative oxide analyses

in the melt. In addition, breccia xenoliths that occur in the Little Ascutney ring-dike formed by explosive eruption (Schneiderman, 1989) and are evidence of the high activity of volatile constituents in the magma. Although the syenite stock intrudes calcareous metasediments, there is no evidence for high $a_{\rm CO_2}$ in the magma whereas the presence of late stage amphibole and mica in the syenite suggests high $a_{\rm Ho0}$.

DISCUSSION

In its lithology, structure, and mineral chemistry, the Ascutney complex resembles other White Magma plutonic-volcanic series complexes: it consists primarily of gabbro-diorite and syenite stocks as well as biotite granite. Like other White Mountain plutonic-volcanic series complexes, Ascutney is calc-alkaline. In addition, a ring dike, so common among other members of this igneous series, occurs as part of the complex.

Behavior of the felsic constituents of the Ascutney intrusions can be modeled in the system Ab-An-Or-Q-H₂O. In the ternary subsystem Ab-An-Or (Fig. 7), bulk compositions plot in the primary phase field of plagioclase indicating that the parent liquids of the intrusions changed compositions by first crystallizing plagioclase which became increasingly sodic. Compositions of plagioclase cores and rims indicate that the liquid compositions moved toward the two feldspar boundary and would have crossed it at the reaction (odd) portion of the boundary curve. Textural evidence for this is perthite rims on plagioclase and the absence of plagioclase in micrographic intergrowths. These features indicate that plagioclase stopped crystallizing soon after alkali feldspar but before quartz appeared; that is, plagioclase ceased growing when the liquids associated with each of the intrusions passed from the plagioclase field, across the two-feldspar boundary into the primary phase field for alkali feldspar (Tuttle and Bowen, 1958).

Variations in mineral chemistry indicate crystallization histories that involved fractionation of early-formed mafic constituents and relative alkali enrichment: unzoned augite has rare orthopyroxene inclusions in the gabbro and coexists with hornblende and biotite in the diorite; zoned amphibole in the gabbro-diorite decreases in Al, Ti, and alkalis from core to rim; though uncommon, clinopyroxene occurs as inclusions in biotite and hornblende in the syenite and the granite; plagioclase is normally zoned from calcic cores to sodic rims in all intrusions. Early crystallization of orthopyroxene, augite, and calcic plagioclase depleted in the magmas in MgO and CaO while crystallization of Fe-Ti oxides depleted them in FeO and Fe₂O₃.

Isotopic data of Foland et al. (1985, 1988) indicate that the gabbro-diorite rocks of the Ascutney complex were derived by fractional crystallization of a mantle-derived basaltic magma that assimilated variable amounts of country rock. Their data indicate that the granite magma evolved, by fractional crystallization with no crustal contamination, from the mafic magma that formed the gabbro-diorite.

The major element data of this study, however, do not



Fig. 5. Photomicrograph showing annite + quartz rimming fayalite + potassium feldspar in quartz syenite. Ascutney Mountain, Vermont, cross-polarized light (sample 5234).

unequivocally support the conclusions of Foland et al. (1985, 1988). Though the gabbro-diorite has high normative An in comparison to the granite (Fig. 7) which would suggest extensive fractional crystallization in deriving the granite from the gabbro-diorite, the Mg numbers of the hornblende in the gabbro-diorite and granite are nearly identical and imply minimal fractional crystallization.

Foland et al. (1985, 1988) did not consider the evolution or parentage of the magma which formed the syenite. Petrography and major element data from this study indicate that the syenite magma was not related to the granite magma. Feldspars in the granite have sharp grain boundaries and are unaltered, unlike those in the syenite



Fig. 6. Stability curves for mafic silicates in Ascutney quartz syenite (after Eugster and Wones, 1962; Thomas, 1982). Also plotted is the granite minimum melting curve (Tuttle and Bowen, 1958). Shaded area indicates P-T region where annite and hastingsite are stable in a granitic melt.



Fig. 7. Bulk compositions of gabbro-diorite and granite in terms of normative An-Ab-Or. The diagram shows the approximate liquidus surface for 500 bars. The approximate position of the quartz- and H₂O-saturated two-feldspar boundary curve is shown. Whole-rock analyses of Daly (1903) and Foland et al. (1985).

which have irregular grain boundaries and are intergrown with quartz, indicating relative H_2O enrichment. Both the amphibole and biotite in the syenite have lower Mg numbers than do those minerals in the granite. Also, amphibole in the syenite has substantial ^[4]Al. In contrast, the normative proportions of An-Ab-Or and the mineral compositions in the gabbro-diorite and syenite may be consistent with a distant relation between these two parent magmas.

SUMMARY AND CONCLUSIONS

Early igneous activity at Ascutney Mountain involved emplacement of radial mafic dikes and glassy felsite dikes with spherules (Balk and Krieger, 1936). Phreatomagmatic explosions occurred, perhaps as a result of the low total pressure and high $a_{\rm H_2O}$, and resulted in the formation of maar-type breccias (Schneiderman, 1989). A volcanic edifice formed on the Cretaceous erosion surface, the remnants of which are preserved as a screen and inclusions in Ascutney syenite.

The three stocks and ring dike of the Ascutney complex reached crustal levels as high as the volcanic edifice itself. Intrusion of gabbro-diorite was followed by intrusion of a syenite ring dike and stock. The gabbro-diorite was cool prior to intrusion of the syenite ring dike as indicated by fine-grained porphyritic texture and chilled margins in syenite against the gabbro-diorite. The syenite stoped fragments of the surrounding country rock and intruded the volcanic structure; xenoliths of country rock and volcanics are contained in it. The last phase of igneous activity was the intrusion of biotite granite into the syenite stock. This study shows that rocks of the Ascutney Mountain igneous complex were emplaced into Precambrian and Paleozoic country rocks at low pressure (~ 2 kbar) and moderate temperature (890–1000 °C) with f_{o_2} below the FMQ buffer. H₂O activity, calculated from biotite compositions in the sygnite was high.

Trends in the compositions of mafic silicates and feldspars in the gabbro-diorite, syenite, and granite suggest that the magmas which formed the gabbro-diorite and granite were possibly unrelated whereas there may have been a distant relation between the gabbro-diorite and syenite magmas.

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