

## Sabelliite, a new copper-zinc arsenate-antimonate mineral from Sardinia, Italy

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**Abstract:** Sabelliite, ideally  $(\text{Cu,Zn})_2\text{Zn}[(\text{As,Sb})\text{O}_4](\text{OH})_3$ , is a new mineral from Is Murvonis Mine, Domusnovas, Sardinia. It occurs as intergrown aggregates or isolated crystals in a compact quartzitic matrix, in an outcrop close to worked fluorite veins of an old mine. The platy {001} disk-shaped emerald-green crystals are transparent with adamantine lustre. Associated minerals are theisite, malachite, azurite, and tetrahedrite. Electron microprobe analyses give: Cu 34.80, Zn 15.65, Si 0.33, As 15.04, Sb 4.96, O 28.00, H 0.67 (calculated), total 99.45 wt%. Optically, sabelliite is nonpleochroic, uniaxial negative, with  $\omega = 1.802(2)$  and  $\epsilon = 1.797(2)$ .  $\text{VHN}_{15}$  microhardness = 386. Density (calc.) =  $4.65 \text{ g/cm}^3$ . The unit cell is trigonal with  $a = 8.201(1)$ ,  $c = 7.315(1) \text{ \AA}$ ,  $V = 426.07(9) \text{ \AA}^3$ , space group  $P\bar{3}$ , and  $Z = 3$ . The strongest lines in the Gandolfi X-ray pattern are ( $d_{\text{obs}}$ ,  $l, hkl$ ) 2.522(100)(121), 2.166(88)(122), 1.805(92)(123), 1.550(100)(410), 1.513(85)(124). The IR spectrum is also given. The mineral is named after Dr. Cesare Sabelli, researcher of the C.N.R.. Sabelliite belongs to a proposed mineral family that includes spangolite, schulenbergite, namuwite, and theisite.

**Key-words:** sabelliite, new mineral, copper-zinc arsenate-antimonate, Is Murvonis Mine (Sardinia, Italy).

### Introduction

In 1990, during field work in Southern Sardinia for an ongoing study of copper alteration minerals, a few specimens were found of an unknown mineral. A preliminary investigation of the material, which was collected in the abandoned Is Murvonis Mine (Domusnovas, Iglesias district), suggested the mineral might be a new species (Brizzi *et al.*, 1993). Shortly thereafter, a mineral showing similar morphological and chemical characteristics was found at mine dumps near Schwaz, Austria (Dillen, 1993, pers. comm.).

A full mineralogical study confirms that the Is Murvonis mineral really is a new species. We are pleased to name the new mineral sabelliite in honour of Dr. Cesare Sabelli, of the Consiglio Nazionale delle Ricerche (C.N.R.), in recognition

of his important work with the copper alteration minerals of Sardinia. The new mineral and name have been approved by the I.M.A. Commission on New Minerals and Mineral Names. A type specimen is kept at the Museum of Natural History of the University of Florence, under catalogue number 2053/RI.

The crystal structure of sabelliite was determined by Olmi *et al.* (1995).

### Occurrence

Geologically, the Iglesias region (S-W Sardinia) consists of a Cambrian-Silurian succession (Carmignani *et al.*, 1986). The Domusnovas area offers a good example of the geology of the region; the main formations comprise a limestone known as the "Calcere Ceroide" as well as poly-

genic Ordovician conglomerates and coarse sandstones with interbedded shales. During Hercynian deformation, the area underwent low-grade metamorphism with the formation of a complex system of faults. Karst phenomena later developed, and hydrothermal fluids subsequently deposited fluorite vein-like bodies along the faults and karst rims occurring in the dark-gray limestone. Only one body of economic interest has been mined in this century. The green or blue fluorite always contains traces of copper (Uras, 1958).

The fluorite is frequently accompanied by sulphides and sulphosalts (galena, sphalerite, chalcopyrite, tetrahedrite-tennantite), and by copper-zinc alteration minerals (aurichalcite, malachite, azurite, brochantite, rosasite, hemimorphite). Sabelliite occurs in a quartzitic outcrop as rare minute crystals often overgrown on foliated aggregates of thesitate, to which it is closely related from a chemical and physical standpoint. Both minerals are believed to be weathering products of sulphides and sulphosalts. Sabelliite is probably the last-formed mineral of the oxidation process.

### Physical properties

Sabelliite has a highly characteristic appearance: its crystals are tabular {001} and cylindrical, thus resembling small coins (no crystal form other than the basal pinacoid has been recognized) (Fig. 1 and 2). The diameter of intergrown disk-shaped crystals rarely exceeds 400  $\mu\text{m}$ , while isolated single crystals can be up to 200  $\mu\text{m}$  in diameter and 15  $\mu\text{m}$  in thickness. They are

transparent and emerald-green, having an adamantine lustre and a light-green streak. The crystals are brittle, with no cleavage, and have a micro-hardness of VHN = 386 (average of 20 determinations with 15 g indenter: range 322-469), which corresponds to a calculated Mohs' hardness of about 4.5.

The mineral's density is too high to be measured using floatation techniques, and not enough material was available for the pycnometer (density bottle) method. The calculated density is 4.65  $\text{g}/\text{cm}^3$ , on the basis of the experimental chemical composition and assuming three formula units in the unit-cell. The mineral does not fluoresce under long- or short-wave ultraviolet radiation.

Sabelliite is optically nonpleochroic, uniaxial negative, with  $\omega = 1.802(2)$  and  $\epsilon = 1.797(2)$ . Refractive indices were measured in Na gel-filtered light by immersion in mixtures of Cargille oils and sulphur saturated methylene iodide. Since the high values of the indices prevented the use of an Abbé refractometer, the indices of the mixtures were determined on an Eickhorst Modul 1 gemmological refractometer (Hurlbut, 1984).

Calculation of the Gladstone-Dale relationship gives  $K_p = 0.1720$  and  $K_c = 0.1753$ , which indicate a "superior" compatibility index  $1-(K_p/K_c)$  of 0.019 (Mandarino, 1981).

### Chemistry and infrared study

Sabelliite was analysed with a JEOL JXA-8600 electron microprobe (15 kV accelerating voltage, 10 nA beam current- Faraday cup, 5  $\mu\text{m}$  defo-

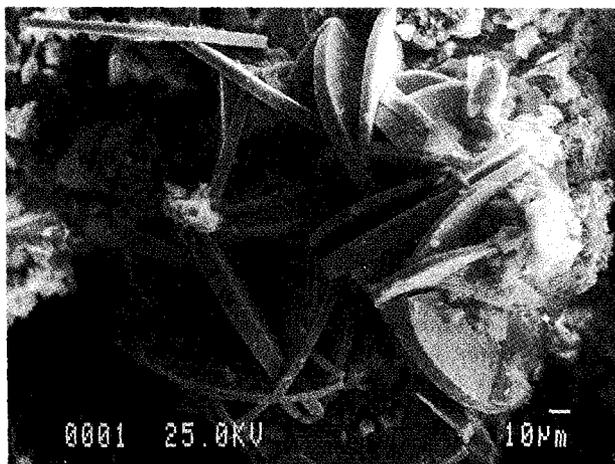


Fig. 1. Electron micrograph of some disk-shaped sabelliite crystals.

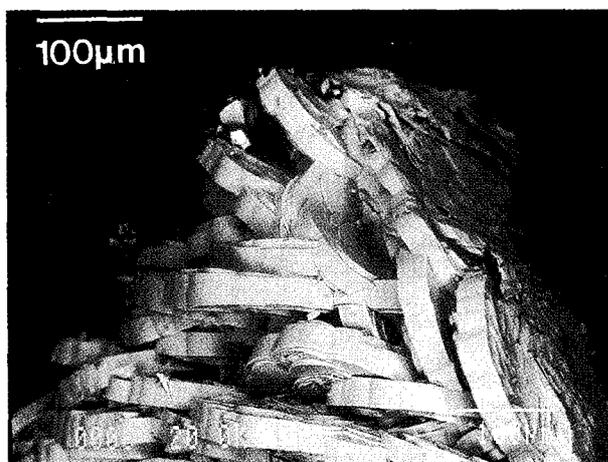


Fig. 2. Electron micrograph of an aggregate of strongly intergrown sabelliite crystals.

cused beam) equipped with a Tracor-Northern Series II system. No elements (with  $Z > 9$ ) other than those reported here were detected by a preliminary 300s-energy-dispersive scan. The standards used were: cuprite ( $\text{CuK}\alpha$ ), willemite ( $\text{ZnK}\alpha$ ,  $\text{SiK}\alpha$ ), arsenolite ( $\text{AsL}\alpha$ ) and senarmonite ( $\text{SbL}\alpha$ ). The dearth of pure material prevented the quantitative determination of the  $\text{H}_2\text{O}$  content. For this reason, the oxygen content was

also measured by electron microprobe with a LDE1 crystal, which is a layered synthetic dispersion lattice (Nash, 1992) manufactured by the Jeol Company, using quartz as standard (McGuire *et al.*, 1992). The raw data were reduced using a  $\phi(\rho z)$  correction routine (Bastin *et al.*, 1987). The results of the analyses (twenty different spots on four crystals) are given in Table 1. The empirical formula, calculated on

Table 1. Electron microprobe analytical results.

|    | 1                   | 2                       | 3            |
|----|---------------------|-------------------------|--------------|
| Cu | 34.80 (32.44-37.26) | CuO                     | 43.56 40.61  |
| Zn | 15.65 (13.91-17.34) | ZnO                     | 19.48 20.77  |
| Si | 0.33 (0.19-0.49)    | $\text{SiO}_2$          | 0.71 -       |
| As | 15.04 (13.42-16.32) | $\text{As}_2\text{O}_5$ | 23.07 23.46  |
| Sb | 4.96 (4.10-6.15)    | $\text{Sb}_2\text{O}_5$ | 6.59 8.26    |
| O  | 28.00 (26.94-29.20) | $\text{H}_2\text{O}$    | 6.05 6.90    |
|    | 98.78               |                         | 99.46 100.00 |
| H  | 0.67                |                         |              |
|    | 99.45               |                         |              |

1 - Atomic wt.% (ranges in parenthesis), including the O content, experimentally determined.  
The 0.67 wt% H amount was calculated to compensate the excess of negative charge.  
2 - Converted oxide percentages.  
3 - Calculated oxide percentages for the idealized formula, with  $\text{Cu}/\text{Zn}=2$  and  $\text{As}/\text{Sb}=4$ .

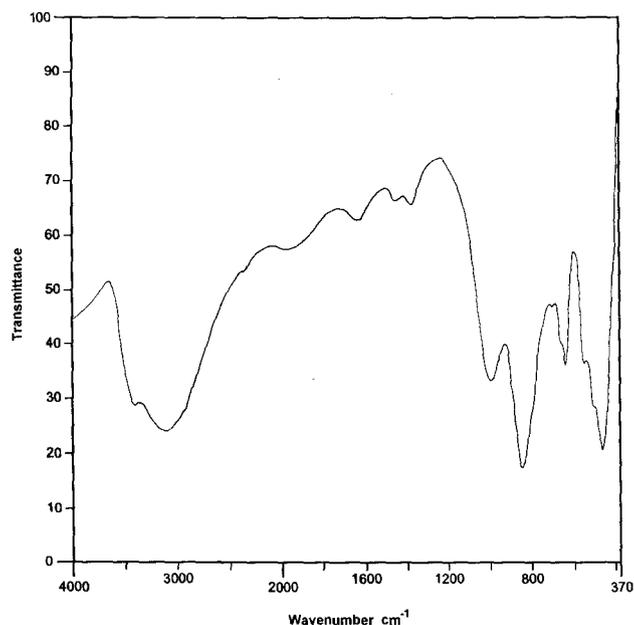


Fig. 3. Infrared absorption spectrum of sabelliite.

the basis of  $O = 7$ , is  $Cu_{2.19}Zn_{0.96}(As_{0.80}Sb_{0.16}Si_{0.05})\Sigma_{1.01}O_{4.31}(OH)_{2.69}$ . The crystal structure refinement (Olmi *et al.*, 1995) shows that the Zn atoms are coordinated both octahedrally (as are the Cu atoms) and tetrahedrally. Further occurrences of sabelliite may display some variation in cationic content. Owing to this possibility, the simplified formula  $(Cu,Zn)_2Zn[(As,Sb)O_4](OH)_3$  was adopted.

For the IR study, a very small amount of material was examined with a FTIR Perkin-Elmer SYS 2000 spectrophotometer. Several IR spectra (which were reproducible to  $\pm 4$   $cm^{-1}$  in the 4000–2000  $cm^{-1}$  region, and to  $\pm 2$   $cm^{-1}$  in the 2000–370  $cm^{-1}$  region) were recorded in the 4000–370  $cm^{-1}$  range, using a KBr micropellet. The spectrum (Fig. 3) confirms the absence of structural  $H_2O$  (Olmi *et al.*, 1995), as there is simply a broad intense band in the 3500–3000  $cm^{-1}$  region. The weak peak at 1636  $cm^{-1}$  may be attributed to water adsorbed by the KBr micropellet (Braithwaite, 1983), or to a harmonic of lower-frequency fundamentals (Sumin de Portilla, 1974; Hill, 1976). As discussed in the structural study, the presence of one ascertained OH group and of an additional disordered hydroxyl appears confirmed by the two bands at 3420 and

3126  $cm^{-1}$ , which are within the region of the stretching vibrations of the hydroxyl groups (Braithwaite & Ryback, 1988). The spectrum is more complex in the region with frequencies lower than 1000  $cm^{-1}$ . The strong absorption bands at 844 and 457  $cm^{-1}$ , with shoulders at 795 and 504  $cm^{-1}$ , are due to the stretching and bending vibrations of the  $AsO_4^{3-}$  group (Farmer, 1974). The absorption at 993  $cm^{-1}$  is probably due to the deformation of OH groups bonded to Cu atoms (Braithwaite, 1983; Sumin de Portilla, 1974), while the absorption at 545  $cm^{-1}$  can be related to stretching vibrations of Zn–O and Cu–O bonds (Tarte & Preudhomme, 1970; Sumin de Portilla, 1974; Sumin de Portilla *et al.*, 1981; Braithwaite, 1983 and 1987).

### X-ray crystallography and related species

Single-crystal X-ray studies (Weissenberg camera and single-crystal diffractometer) show that sabelliite is trigonal. The space group, from the crystal structure analysis, is  $P\bar{3}$ .

The X-ray powder diffraction data given in

Table 2. X-ray powder diffraction pattern for sabelliite (114.6 mm Gandolphi camera, Ni-filtered CuK $\alpha$  radiation, photometrically measured intensities).

| <i>hkl</i> | <i>d</i> <sub>calc.</sub><br>(Å) | <i>d</i> <sub>obs.</sub><br>(Å) | <i>I</i> / <i>I</i> <sub>0</sub> | <i>hkl</i> | <i>d</i> <sub>calc.</sub><br>(Å) | <i>d</i> <sub>obs.</sub><br>(Å) | <i>I</i> / <i>I</i> <sub>0</sub> |
|------------|----------------------------------|---------------------------------|----------------------------------|------------|----------------------------------|---------------------------------|----------------------------------|
| 001        | 7.314                            | 7.33                            | 42                               | 412        | 1.427                            | 1.427                           | 54                               |
| 100        | 7.102                            | 7.15                            | 7                                | 224        | 1.365                            | 1.364                           | 12                               |
| 101        | 5.096                            | 5.11                            | 18                               | 025        | 1.353                            | 1.352                           | 8                                |
| 110        | 4.101                            | 4.11                            | 55                               | 240        | 1.342                            | 1.341                           | 23                               |
| 002        | 3.657                            | 3.66                            | 52                               | 241        | 1.320                            | 1.320                           | 55                               |
| 111        | 3.577                            | 3.58                            | 49                               | 413        | 1.308                            | 1.308                           | 13                               |
| 102        | 3.252                            | 3.25                            | 31                               | 125        | 1.285                            | 1.284                           | 25                               |
| 201        | 3.195                            | 3.20                            | 46                               | 242        | 1.260                            | 1.260                           | 16                               |
| 112        | 2.729                            | 2.726                           | 29                               | 512        | 1.204                            | 1.204                           | 9                                |
| 120        | 2.684                            | 2.688                           | 29                               | 414        | 1.182                            | 1.183                           | 22                               |
| 022        | 2.548                            | 2.548                           | 15                               | 243        | 1.176                            | } 1.175                         | 31                               |
| 121        | 2.520                            | 2.522                           | 100                              | 315        | 1.175                            |                                 |                                  |
| 030        | 2.367                            | 2.369                           | 26                               | 405        | 1.129                            | 1.129                           | 8                                |
| 031        | 2.252                            | 2.253                           | 28                               | 216        | 1.110                            | 1.109                           | 23                               |
| 122        | 2.164                            | 2.166                           | 88                               | 244        | 1.082                            | 1.082                           | 26                               |
| 113        | 2.096                            | 2.096                           | 25                               | 415        | 1.064                            | 1.064                           | 24                               |
| 220        | 2.050                            | 2.054                           | 12                               | 440        | 1.025                            | 1.025                           | 7                                |
| 221        | 1.974                            | 1.974                           | 18                               | 530        | 1.015                            | } 1.014                         | 17                               |
| 311        | 1.902                            | 1.901                           | 18                               | 070        | 1.015                            |                                 |                                  |
| 004        | 1.829                            | 1.828                           | 7                                | 531        | 1.005                            | } 1.005                         | 31                               |
| 123        | 1.805                            | 1.805                           | 92                               | 701        | 1.005                            |                                 |                                  |
| 132        | 1.734                            | 1.733                           | 23                               | 245        | 0.989                            | 0.989                           | 12                               |
| 041        | 1.725                            | 1.725                           | 23                               | 532        | 0.978                            | } 0.978                         | 18                               |
| 114        | 1.670                            | 1.670                           | 26                               | 702        | 0.978                            |                                 |                                  |
| 204        | 1.626                            | 1.626                           | 18                               | 703        | 0.937                            | } 0.937                         | 23                               |
| 321        | 1.590                            | 1.591                           | 22                               | 533        | 0.937                            |                                 |                                  |
| 223        | 1.569                            | 1.569                           | 9                                | 360        | 0.895                            | 0.895                           | 15                               |
| 410        | 1.550                            | 1.550                           | 100                              | 361        | 0.888                            | } 0.888                         | 27                               |
| 313        | 1.532                            | 1.532                           | 8                                | 704        | 0.887                            |                                 |                                  |
| 124        | 1.511                            | 1.513                           | 85                               | 534        | 0.887                            |                                 |                                  |
| 322        | 1.488                            | 1.489                           | 17                               | 705        | 0.834                            | } 0.834                         | 15                               |
| 005        | 1.463                            | 1.462                           | 11                               | 535        | 0.834                            |                                 |                                  |
| 034        | 1.447                            | 1.448                           | 16                               |            |                                  |                                 |                                  |

Table 2 were obtained using a Gandolphi camera (114.6 mm diameter) with Ni-filtered CuK $\alpha$  radiation. Indexing was performed on the basis of single-crystal data, taking account of the intensities of the reflections collected for the structural study. The unit-cell parameters, refined by means of a least-squares calculation from the powder data, are:  $a = 8.201(1)$ ,  $c = 7.315(1)$  Å,  $V = 426.07(9)$  Å<sup>3</sup>.

The chemical composition, unit-cell parameters and symmetry of sabelliite are fairly similar

to those of several other copper minerals, which are listed in Table 3. The powder patterns of all these species also show significant similarities. The lattice parameters of sabelliite are practically identical to those of schulenbergite (Hodenberg *et al.*, 1984). The chemical variability assumed for sabelliite is actually observed in schulenbergite: its Zn-analogue, indeed, has a highly variable Cu/Zn ratio, whose lowest value is 0.34 (Livingstone *et al.*, 1992), compared with the value of 1.33 found in schulenbergite. Spangolite

Table 3. Formulae and lattice parameters of related minerals.

| Mineral        | Chemical formula  | Cell parameters (Å)  | Symmetry           |
|----------------|---|----------------------|--------------------|
| Spangolite     | Cu <sub>6</sub> Al(SO <sub>4</sub> )(OH) <sub>12</sub> Cl·3H <sub>2</sub> O                                 | $a=8.274$ $c=14.371$ | $P31c$             |
| Namuwite       | (Zn,Cu) <sub>4</sub> (SO <sub>4</sub> )(OH) <sub>6</sub> ·4H <sub>2</sub> O                                 | $a=8.29$ $c=10.50$   | hexagonal          |
| Schulenbergite | (Cu,Zn) <sub>7</sub> (SO <sub>4</sub> ,CO <sub>3</sub> ) <sub>2</sub> (OH) <sub>10</sub> ·3H <sub>2</sub> O | $a=8.249$ $c=7.183$  | $P3$ or $P\bar{3}$ |
| Theisite       | Cu <sub>5</sub> Zn <sub>5</sub> [(As,Sb)O <sub>4</sub> ] <sub>2</sub> (OH) <sub>14</sub>                    | $a=8.223$ $c=15.019$ | trigonal           |
| Sabelliite     | (Cu,Zn) <sub>2</sub> Zn[(As,Sb)O <sub>4</sub> ](OH) <sub>3</sub>  | $a=8.201$ $c=7.315$  | $P\bar{3}$         |

(Merlino *et al.*, 1992), whose *a* parameter is the same and *c* parameter is twice that of sabelliite or schulenbergite, exhibits dense octahedral sheets, as does sabelliite (Olmi *et al.*, 1995). Namuwite (Bevins *et al.*, 1982) and theisite (Bonazzi & Olmi, 1989), whose structures are still unknown, show the same *a* parameter and a closely similar symmetry.

From the above considerations, and from the structural analogies discussed by Olmi *et al.* (1995), the grouping of these species in the same mineral family appears quite reasonable, following also the suggestion by Hodenberg *et al.* (1984) who have pointed out that schulenbergite is closely related to spangolite and namuwite.

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